

Steady or Sudden: the 2008 Eruption of Okmok, Alaska

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Abstract

The 2008 eruption of Okmok (OK08) had among the shortest run-ups for an arc volcano (~5 hours of heightened seismicity), yet also experienced over a decade of inflation since its prior eruption in 1997. Was OK08 driven by steady inflation or a sudden diking event? To answer this question, we investigated OK08 melt inclusions and olivines to constrain the depth and timing of magmatic events. We homogenized olivine-hosted melt inclusions and found CO₂ concentrations up to 1800 ppm, providing new evidence for a mid-crustal magma storage region (~15 km) beneath Okmok that aligns closely with the deepest earthquakes found during the five-hour precursory seismicity. These high CO₂ melt inclusions are hosted in olivines that are unzoned in Fe-Mg, constraining residence in host magma to less than a day. Hence, the precursory seismicity was magmatic, and the eruption expelled crystals that could have risen five hours prior to eruption. Fe-Mg diffusion timescales and crystal depths also show magma movement beneath Okmok during its inter-eruptive period, with three dominant reservoirs at depths from 15 to 1 km. Mixing took place at least five years prior to eruption with an increasing number of olivines reflecting mixing timescales in early 2008. This magmatic ramp-up aligns with the shifting of the deformation source and proposed tensile failure of the reservoir in the months prior to eruption. The occurrence of

25 both steady and sudden processes culminated in OK08. Sudden explosive eruptions could be a
26 consequence of a magmatic system ‘primed’ by years of slow magma input.

27 **1. Introduction**

28 Volcanic eruptions are often preceded by escalating levels of unrest, known as run-up
29 (Passarelli & Brodsky, 2012). Identifying run-up is critical to our ability to forecast eruptions and
30 issue effective warnings (Pesicek et al., 2018). Extremely short run-ups are particularly
31 problematic for providing timely warnings. When such eruptions occur, it is critical to analyze in
32 hindsight the pre-eruptive geophysical signals to better prepare for future eruptions. The 2008
33 eruption of Okmok volcano (OK08), Alaska, is an example of an eruption with an extremely short
34 run-up, with a mere five hours of heightened seismicity preceding the VEI 4 eruption (Johnson et
35 al., 2010; Larsen et al., 2009; Fig 1a.). In the compilation of Passarelli and Brodsky (2012), OK08
36 is the arc volcano with the shortest run-up globally. Since then, other eruptions have been identified
37 with no apparent run-up (e.g., Pavlof 2016; Pesicek et al., 2018). The precursory behavior of OK08
38 was in stark contrast to the inter-eruptive period (1997–2008), when there was episodic inflation
39 for over a decade but with little seismicity (Freymueller & Kaufman, 2010; Johnson et al., 2010;
40 Lu et al., 2010; Power et al., 2019; Fig. 1b). Approximately two-to-three months before the
41 eruption, there was a distinct shift in the long-term deformation source or change in source
42 geometry (Freymueller & Kaufman, 2010).

43 Recent studies by Albright et al. (2019) and Garza-Girón et al. (2023) propose two
44 opposing interpretations to the trigger of OK08 based on geodetic and seismic data, respectively.
45 Albright et al. (2019) propose Okmok reached critical tensile failure in the months prior to
46 eruption, primed by years of inflation. We refer to this run-up as slow and steady. On the other
47 hand, Garza-Girón et al. (2023)’s used their enhanced earthquake catalogue to suggest a deep and

48 rapid trigger initiating 5 hours before eruption, linked to the opening of a planar feature that
49 extended to the maximum depth of seismicity of ~15 km. This view implicates a sudden run-up to
50 eruption.

51 Okmok thus displayed unrest over many timescales: from years of episodic inflation, to
52 months of change in the inflation source and stress state, to hours in an earthquake swarm that
53 immediately preceded eruption. Such protracted episodes and manifestations of unrest have been
54 termed “ultimate,” “proximal,” and “immediate” by Putirka (2017), and have obvious implications
55 for successful eruption forecasting. Seismicity and deformation are non-unique to specific
56 magmatic processes (McNutt & Roman, 2015; Wauthier et al., 2016), and understanding the
57 warning signs to eruption requires an understanding of the magmatic system – its distribution
58 within the crust, the timing of magma ascent and storage, and factors that control magma
59 buoyancy, viscosity, volatile concentration and composition. Such information can only be
60 provided by studying eruptive products, as they encode in their chemical composition such vital
61 information. Specifically, petrological tools such as diffusion chronometry and solubility
62 barometry can provide a timeseries of storage depth and composition that can then be mapped to
63 geophysical timeseries (Kahl et al., 2022; Lynn et al., 2024; Rasmussen et al., 2018).

64 To explore this well-posed example of steady vs. sudden run-up, we examine here the
65 earliest erupted products of OK08 (Larsen et al., 2013) using 1) olivine diffusion chronometry to
66 provide the timing of magma mixing, 2) volatile solubility to provide the depth of magma residence
67 leading up to OK08 and 3) major and trace element compositions of whole rock, glass and mineral
68 phases to provide constraints on the different magma bodies tapped. With this dataset, we address
69 the following fundamental questions. Is the precursory seismicity related to magma movement?
70 Can we associate syn-eruptive and inter-eruptive geophysical observations with magmatic events?

71 **2. Background**

72 On July 12, 2008, Okmok volcano produced a 16 km eruptive column and extensive tephra
73 fall-out over a period of 4 hours (Larsen et al., 2009, 2015; Unema et al., 2016). Weak ash
74 emissions continued from multiple vents until activity ceased on August 19, 2008. In this study
75 we focus on the July 12 phase of the eruption, which produced scoria fall units deposited inside
76 and outside the caldera (0.048 km³ DRE erupted volume and 9.85 x 10⁶ kg/s discharge rate; Larsen
77 et al., 2015; Unema et al., 2016).

78 OK08 had several unique features. First, the vent shifted away from Cone A, the eruption
79 site over the past two centuries, to form a new cone, Ahmanilix (Larsen et al., 2015; Fig. 1a).
80 Second, OK08 (VEI 4) was the most explosive eruption since Okmok-II, the caldera-forming
81 eruption in BCE43 (Larsen et al., 2013). Third, the geochemistry of the newly erupted products
82 was diverse, spanning from basaltic to andesitic compositions (Larsen et al., 2013). The first ash
83 erupted during OK08 contained a unique population of primitive olivine-hosted melt inclusions
84 (MIs) that were found in neither the 1997 scoria (Zimmer et al., 2010) nor the lapilli/bomb samples
85 of the main July 12 phase (described as scoria in Larsen et al., 2013). Additionally, the whole rock
86 composition of the main July 12 phase was more evolved (55.5 wt.% SiO₂) than the previous 1997
87 and 1958 eruptions (52 wt.% SiO₂; Finney et al., 2008; Larsen et al., 2013). Finally, the temporal
88 pattern of precursory seismicity, predominantly volcano-tectonic (Garza-Giron et al., 2023), did
89 not progress from shallow to deep, unlike the “top-down” behavior of other Aleutian volcanoes
90 (Roman & Cashman, 2018).

91 The current model for OK08 is that the five-hour seismicity was related to dike intrusion
92 from the main storage region (at ~3.4 km below the surface, the depth of the long term deformation
93 source, Lu et al., 2010; Wang et al., 2021) to a shallower, older and chemically evolved body

94 beneath Cone D (Larsen et al., 2013; Ohlendorf et al., 2014; Miller et al., 2020 Fig. 1a). The
95 expulsion of resident magma was supported by the volume change during the eruption (0.14 ± 0.01
96 km^3), $\sim 3x$ greater than the 1997-2008 inflation ($\sim 0.05 \text{ km}^3$) (Lu & Dzurisin, 2010; Fig. 1b). A
97 shallow sill at $\sim 1.4 \text{ km}$ was later identified in the post-2008 inflation (Xue et al., 2020). Seismic
98 tomography studies show a more complex magmatic architecture, including a deeper sill at 2 to 6
99 km and a deeper storage region up to 10 km depth (Kasatkina et al., 2022; Miller et al., 2020;
100 Ohlendorf et al., 2014). Finally, Walwer et al. (2019) proposed that oscillatory deformation
101 following inflation events during the inter-eruptive period could represent a hydraulic connection
102 between two shallow reservoirs. These proposed interactions between deep and shallow magma
103 reservoirs in the decade prior to OK08 motivates this petrological study to constrain the timing
104 and depth of magma mixing beneath Okmok.

105 **3. Materials and Methods**

106 Critical to this study was the distal collection of ash samples erupted in the first few hours
107 of OK08, as this material contained crystals and melt not present later in the July 12 main phase
108 (Larsen et al., 2013). We focus here on the ash sample 08LKOK001 (described in Larsen et al.,
109 2013) and collected on a boat by rancher Lonnie Kennedy while evacuating from Umnak Island.
110 Loose olivine crystals were picked and examined for MIs that were glassy, free of crystals, fully
111 enclosed in olivine, and without cracks, following Rose-Koga et al. (2021) (Fig. 2b and
112 Supplementary Information, SI, -1). We also selected free-standing matrix glasses (MGs) that
113 exhibited minimal microlite growth. For diffusion chronometry, we selected euhedral olivines with
114 adhered matrix glass as an indication of the rim. See SI-1 to SI-7 for detailed methods and
115 Supplementary Dataset, SD, -1 and-2 for all data.

116

117 *3.1 Melt inclusions, matrix glass and bulk ash analyses*

118 To account for CO₂ in MIs that contained vapor bubbles, we homogenized MIs using the
119 piston cylinder method of Rasmussen et al. (2020) at the Lamont-Doherty Earth Observatory
120 (LDEO) (SI-2 and Fig. 2b and 2c). Subsequently, homogenized, bubble-free (untreated) and
121 bubble-bearing (untreated) MIs, and free-standing MGs were doubly polished (Fig. 2c) to quantify
122 H₂O and CO₂ concentrations using a Thermo-Nicolet iN10 Fourier Transform Infrared
123 Spectrometer (FTIR) at LDEO. FTIR spectra (Fig. 2a) were processed using the PyIRoGlass
124 package to account for baseline uncertainties and compositional effects on absorption coefficients
125 (Shi et al., 2024; SI-3). Major elements, sulfur and chlorine in MIs and MGs, and major elements
126 of host olivines were measured on a Cameca SX-5 electron probe micro analyzer (EPMA) at the
127 American Museum of Natural History (AMNH) (SI-4). We also analyzed the major elements of
128 11 MGs adhered to the olivine. A subset of the MIs and MGs were further analyzed for trace
129 elements on a laser ablation inductively coupled plasma mass spectrometer (LA-ICP-MS) at
130 Cornell University (SI-5). MIs were corrected for post-entrapment modifications (Danyushevsky
131 et al., 2000) using MIMiC (Rasmussen et al., 2020) with Fe³⁺/Fe^T fixed at 0.20 (as for 1997 scoria,
132 Andrys, 2023) and using the Fe-Mg (ol/liq) partition coefficient expression in Toplis (2005). The
133 bulk geochemical composition of the initial ash sample, 08LKOK001, was carried out following
134 LiBO₂ flux fusion using an ICP-OES at LDEO (SI-6).

135

136 *3.2 Diffusion chronometry*

137 Olivines were analyzed by backscatter electron imaging using a JEOL JCM-7000
138 NeoScope Scanning Electron Microscope at LDEO to identify different zonation patterns.
139 Chemical profiles were measured using a JEOL-8530F EPMA at the University of Alaska,

140 Fairbanks Advanced Instrumentation Laboratory and at the AMNH with a high sensitivity
141 analytical routine for major and trace element profiles (SI-4). Orientations were later verified using
142 electron backscattering diffraction on a Scanning Electron Microscope JSM-7800F at the Earth
143 Observatory of Singapore. Corrections were made to the diffusivities to account for diffusive
144 anisotropy in olivine (Costa & Dungan, 2005).

145 We modeled olivine Fe-Mg profiles using a 1-D finite difference model (see SI-7 for model
146 design, errors and parameters). Additionally, we also performed forward 1-D and 3-D finite
147 difference modeling to estimate the maximum timescale for which unzoned olivine crystals may
148 reside in non-equilibrium melt and the timescales for the loss of olivine core compositions,
149 respectively.

150 **4. Whole rock, melt inclusions and glass chemistry**

151 ***4.1 Homogenization experiments***

152 After the homogenization experiments, 26 MIs were glassy and free from other phases
153 (Fig. 2b), while 6 MIs crystallized small sulfide globules (due to the lowering of fO_2 during the
154 experiment; (Ding et al., 2025; Taracsák et al., 2023). Five MIs also experienced decrepitation
155 during the experiments, indicated by an excess of K_2O (KBr is used as a cushioning matrix). These
156 MI were corrected using the K_2O-SiO_2 trend defined by the rest of the MIs and are distinguished
157 by a different symbol in the figures (SI-8). CO_2 concentrations in the MIs are not affected
158 significantly by these processes, as experimental homogenization can lead to many-fold increases
159 in CO_2 , and decrepitated MIs can also retain high volatile concentrations (Ding et al., 2025).

160 ***4.2 Major and trace elements***

161 The bulk ash from the initial phase of the eruption is more mafic (52.5 wt.% SiO₂) than the
162 main July 12 phase of the eruption (55 wt.% SiO₂) (SD-1). The initial ash is also significantly
163 richer in phenocrysts (33% crystals, SI-9) compared to the largely aphyric main phase scoria
164 (Larsen et al., 2013). Compositions of MIs and MGs analyzed in this study overlap well with those
165 in Larsen et al. (2013) (Fig. 3a). The offset in SiO₂ between the homogenized MIs and untreated
166 MIs studied here is a function of selecting mostly bubble-free MIs for the untreated population,
167 and not related to the homogenization procedure. The bubble-bearing MIs from the initial ash
168 contain low SiO₂ (<50 wt. %) compositions not observed in the bubble-free MI population, nor in
169 any of the later-erupted products, nor in 1997 MIs (Larsen et al., 2013; Zimmer, et al., 2010). Thus,
170 olivines with unique, primitive melts erupted in early OK08 only. The large range in SiO₂ observed
171 in the early MGs reported by Larsen et al. (2013) is due in part to differences between free-standing
172 (55–57 wt. %) and olivine-adjacent MGs (51–54 wt. %). The more mafic melt surrounding olivine
173 grains is not caused by olivine dissolution or boundary layer effects, as incompatible elements
174 such as Ca and Cl do not deviate from the trend defined by all other MIs and MG (Fig. 3a and SI-
175 10; Shea et al., 2022). Thus, the olivine-adjacent MG appears to be a separate liquid that did not
176 mix with the resident, more silicic melt.

177 Figure 4 and S11 show the trace element variations in MIs, MGs and WRs erupted from
178 Okmok since the last caldera-forming eruption in BCE 43. All compositions define a coherent
179 liquid line of descent (LLD) in V, with a clear kink when magnetite appears along the cotectic
180 (and strongly partitions V). The MIs from the initial ash were entrapped before magnetite
181 crystallization, while the free-standing MGs and 2008 main July 12 phase WR formed after
182 magnetite crystallization (Fig. 4b). Cr is singled out by Larsen et al. (2013) as a distinguishing
183 feature of OK08 relative to the adjacent Cone B magmas from the 1817 eruption. Olivine host

184 Forsterite (Fo) content decreases with decreasing Cr in the MI, consistent with co-crystallization
185 of olivine, clinopyroxene and Cr-spinel. The low SiO₂ MIs have elevated Cr contents above any
186 of the post-caldera WRs (Fig. 4a). For Sc, the MIs and MGs fall within the scatter of WRs which
187 rules out melting of pyroxene-rich lithology to explain the low SiO₂ and high CaO MIs (Larsen et
188 al., 2013). Instead, these low SiO₂ - high CaO MIs have elevated Al₂O₃ and positive Eu anomalies,
189 consistent with plagioclase dissolution. Despite the wide variety in major element compositions,
190 all OK08 MI, MG and WR fall on the same Th vs. Nb trend, indicating a common parentage (Fig.
191 4e).

192 *4.3 Volatiles and saturation pressures*

193 Water concentrations span between 0.77–2.4 wt.% for homogenized MIs and 0.92 – 1.75
194 wt.% for bubble-free MIs, within the range found previously in the same sample (Larsen et al.,
195 (2013), with the exception of one MI with an anomalously high concentration of 3.6 wt.% H₂O).
196 The free-standing MGs record lower water contents, 0.5–1.02 wt.% (Fig. 3b).

197 Homogenized MIs that crystallized sulfide have S concentrations between 1000–1200
198 ppm, consistent with the predicted concentration at sulfide saturation (Smythe et al., 2017).
199 Samples without sulfide have S concentrations of 118–1443 ppm for bubble-bearing MIs and 179–
200 532 ppm for bubble-free MIs. S concentrations are below detection limit (~90 ppm) in MGs. The
201 maximum S concentration in OK08 MIs (1443 ppm) is similar to that from other Okmok eruptions
202 (~1300 and ~1600 ppm in 1997 and Okmok II, respectively; Zimmer et al., 2010, Peccia et al.,
203 2023).

204 The majority of MIs and all MG have CO₂ below detection limit (< 100 ppm, calculated
205 using equations in Shi et al. (2024) and dependent on sample thickness). Only 11 of the 32
206 homogenized MIs have detectable CO₂, which ranges from 150 to 1772 ppm (Fig. 3b). All these

207 values are substantially higher than the maximum of 49 ppm CO₂ found in untreated samples
208 (Larsen et al., 2013), confirming the well-known importance of shrinkage bubbles in quantifying
209 the total inventory of CO₂ hosted within MIs (Hartley et al., 2014; Moore et al., 2015). The very
210 low S (<532 ppm), CO₂ concentrations (detection limit, <62 ppm) and evolved compositions in
211 bubble-free MIs (>52 wt.% SiO₂) also calls into question the approach of several studies that only
212 focus on bubble-free MIs (e.g., H el ene et al., 2023; Hlinka et al., 2021). While avoiding the bubble
213 problem, study of such MIs may bias results to the shallow pressures and more evolved
214 compositions.

215 In order to calculate the vapor-saturation pressures of the MIs, we use the python package
216 VESIcal (Iacovino et al., 2021; Wieser et al., 2022) with the solubility model MagmaSat (Ghiorso
217 & Gualda, 2015), as it covers the range of silica observed in MIs and MGs from OK08 (46–57
218 wt.% SiO₂, Fig. S12). MIs record saturation pressures up to 409 MPa (~15 km), with most
219 recording less than 100 MPa (~4 km; Fig. 3b; using the Okmok-specific density model in
220 Rasmussen et al. (2022) to iteratively convert pressures to depths).

221 As the majority of the MIs analyzed here have CO₂ below detection limit and could have
222 experienced diffusive H₂O loss (Barth et al., 2019; Portnyagin et al., 2008) (Fig. 3b and S13), their
223 saturation pressures place them within a shallow storage region (~1–2 km), representing their last
224 depth of re-equilibration. On the other hand, the MI that records the highest pressure (~15 km)
225 provides the first direct evidence for magmatic inputs from the mid-crust feeding Okmok
226 eruptions.

227 **5. Olivine populations and timescales**

228 Olivine phenocrysts range in composition from Fo₇₂ to Fo₈₆ (Fig. 5a), with both zoned and
229 unzoned populations. The decoupling of the zonation profiles across various elements suggests
230 chemical diffusion rather than growth zoning (SI-14). Interestingly, olivine rims show limited
231 variation (Fo₇₇₋₈₂), while cores show the entire range of measured Fo (Fo₇₂₋₈₆), reflecting
232 equilibration with the wide range of OK08 liquids. Fo₈₆ cores are in equilibrium with the most
233 primitive MIs; Fo₇₂ cores are in equilibrium with the free-standing MG and main scoria WR.

234 To further characterize the population of olivine crystals from this eruption, we evaluate
235 the backscatter electron (BSE) images of 72 randomly selected crystals (Fig. 5b). In this way, we
236 reduce bias from crystals analyzed for diffusion chronometry, where unbroken, zoned crystals
237 were preferentially selected. We categorize these diverse olivines into four populations. P₁, sharp
238 profile (Fo plateaus < 100 μm from rim); P₂, shoulder profile; P₃, long profile (Fo plateaus > 100
239 μm from rim). P₄, unzoned (SI-15).

240 ***5.1 Crystal timescales***

241 Most olivine chemical profiles do not display a distinct sigmoidal shape, and thus initial
242 conditions were difficult to determine. Some profiles show enrichment in P, often associated with
243 a rapid growth boundary (Milman-Barris et al., 2008), allowing assignment of an initial step in the
244 compositional profile. In cases where we were unable to distinguish a step or obtain an acceptable
245 fit, we modeled two timescales: one initially homogenous and the other with a step profile (as in
246 Rasmussen et al., 2018), and reported the timescale with the lower error. A recent study by Shea
247 et al. (2023) finds that melt-bearing experiments show enhanced diffusivities in olivine, ~ 10x that
248 in melt-free experiments of DC07 (Dohmen & Chakraborty, 2007). We report in SD-1 results
249 using both the Shea et al. and DC07 diffusivities.

250 Crystal timescales, using DC07 diffusivities, span from weeks to five years before the
251 eruption (Fig. 5c and 6). We binned the timescales into three general populations: 1) weeks (Fig.
252 6a), 2) months (Fig. 6b) and 3) years (Fig. 6c). These bins also help to assign approximate
253 timescales to the BSE profiles: P₁ (sharp) profiles reflect magma mixing within weeks to months
254 before the eruption, while P₃ (long) profiles experienced magma mixing more than a year prior to
255 eruption.

256 With timescales of years before OK08, diffusion may reach olivine cores and thereby
257 confound efforts to use the core composition as the initial condition. Using the measured core
258 composition in these cases would underestimate the true time (Shea et al., 2015; Mourey et al.,
259 2023). To test the extent of loss of initial core composition, we used the 3-D finite difference model
260 by Shea et al. (2015) for an initially homogenous Fo₇₂ crystal surrounded by magma in equilibrium
261 with Fo₈₀ with crystal diameters of 500 μm (mean crystal size in this eruption) and 1 mm (max;
262 Fig. 7f). We extracted profiles along the [001] axis through the core of the crystal along the [100]-
263 [001] plane, which is the most ideal transect for 1-D profiles (Shea et al., 2015). For OK08, 500
264 μm crystals would have lost ~2 mol % Fo after a year of diffusion while 1 mm crystals would still
265 retain initial core compositions (Fig. 7f). Evidence for loss of core compositions is shown in Fig.
266 5c, where the longest (red) timescales do not have end-member core compositions (Fo₇₂ or Fo₈₆).
267 Thus, the Fo diffusion clock may lose resolution after a year and not capture older timescales for
268 crystals with longer residence in the Okmok system (Mourey et al., 2023).

269 Some olivines appear unzoned in their EPMA profiles (Fig. 7d and S10), which generally
270 start within 5 to 10 μm from the rim (Barth et al., 2024; Lynn et al., 2017). Higher resolution BSE
271 images of some of “unzoned” crystals do show a change in greyscale (Fe/Mg) within 10 μm of the
272 rim (SI-16), possibly representing quench overgrowths similar to those observed in Rasmussen et

273 al. (2018). However, if those changes in greyscale represent diffusion, we can estimate the
274 maximum time the crystal could reside in a non-equilibrium magma and remain “unzoned” in
275 EPMA profiles. We use a forward 1-D finite difference model with DC07 diffusivities to obtain
276 expected profiles along the [001] and [100] axes (diffusively fast and slow, respectively). We
277 modeled an initially homogenous end-member crystal, Fo₈₆, mixing with a magma in equilibrium
278 with the mean rim composition Fo₈₀ (Fig. 7d) and consider both axes, as the unzoned crystals were
279 not oriented. Assuming no growth or resorption, this model crystal would show no detectable Fo
280 change beyond 5 microns from the crystal rim within two to three days for the [100] axis and less
281 than a day for the [001] axis. Some of the unzoned olivines were experimentally treated to
282 homogenize their MIs and may have experienced some resorption or diffusion. However, given
283 the distinct euhedral faces (Fig. 7a, b and 7c) and short laboratory heating times (minutes), any
284 resorption and diffusion should not significantly alter the observed profiles. Thus, we can conclude
285 that unzoned crystals that are not in equilibrium with the host melt have a maximum residence
286 time of less than a day.

287 **6. Discussion**

288 ***6.1 Diversity of melt inclusions and architecture of the Okmok system***

289 The OK08 tephra contains a diversity of MIs and olivines that are both more primitive and
290 more evolved in composition than the products of the last eruption in 1997 (Fig. 3a). To explain
291 these diverse compositions, we propose at least three distinct reservoirs were tapped during OK08.
292 The chemical composition and depth of the magma in these reservoirs is constrained by our MI-
293 MG analyses. The first reservoir is what we consider the main reservoir (R_m), coincident with the
294 long-term inflation source beneath Okmok. The composition of R_m is similar in bulk composition
295 to the magmas erupted from Cone A in 1958 and 1997 (52–53 wt.% SiO₂, aphyric). The Cone A

296 WRs are in equilibrium with Fo_{76–82} (assuming Fe³⁺/Fe^T of 0.2 and K_{D ol/melt} = 0.3, Fig. 5a), which
297 overlaps well with the olivine rims in OK08. The MG surrounding the OK08 olivines is also
298 similar in composition to the 1997 WR (Fig. 3, 4). This suggests tapping of the same storage region
299 in 1997 and 2008 (Finney et al., 2008; Larsen et al., 2013; Zimmer et al., 2010). Additionally,
300 Fo_{79–82} olivines host MIs with Cr contents that overlap with the 1958, 1997 and historical Cone A
301 and Cone D WRs (Fig. 4a). These intermediate compositions (51–53 wt.% SiO₂) are not created
302 by the mixing of the primitive and evolved endmembers, as constrained by the kinked V-SiO₂
303 LLD (Fig. 4b, where the 1997 WR plots at the maximum V). Additionally, given that the inflation
304 between 1992–2008 remained consistent in depth and location (Fournier et al., 2009; Lu et al.,
305 2005; Lu & Dzurisin, 2010), except for the 3–4 months prior to OK08 (Freymueller & Kaufman,
306 2010), this liquid in equilibrium with ~Fo_{76–82} likely represents the composition of Okmok’s long
307 term main storage region at ~3 km (Fig. 8d). The MG surrounding OK08 olivines represents the
308 R_m liquid. We note that the composition of the initial ash (52.5 wt.% SiO₂ and Fo₈₀ equilibrium
309 olivine) is also consistent with derivation from R_m and implicates its early tapping.

310 Although R_m is important to the long-term geodetic signal at Okmok and the source of the
311 initial ash, it was not the source of the main July 12 phase of OK08, which we propose was derived
312 from a shallower reservoir, R_s. The main July 12 phase of OK08 consisted of more evolved,
313 uniform and largely aphyric WR compositions with 55–56 wt.% SiO₂, like the average of the free-
314 standing MG in the initial ash, and in equilibrium with the most evolved olivines, Fo_{71–73}, (Fig.
315 5a). Given that this composition is more evolved than the magma erupted in 1997, R_s may represent
316 the resident reservoir beneath Cone D (Larsen et al., 2013). R_s is likely located at 1–2 km depth,
317 as reflected by the common CO₂-and H₂O poor MIs measured in this study (Fig. 8a). These depths
318 are much shallower than the 2–6 km stacked sills identified seismically by Miller et al. (2020) and

319 Kasatkina et al. (2022), but agree with the sill identified by Xue et al. (2020) in the post-2008
320 inflation at around 1.4 km (Fig. 8d).

321 The third reservoir, R_d , represents the source of the deep, mafic (<50 wt.% SiO_2) melt
322 preserved in bubble-bearing MIs from the initial phase of OK08. Some of these MIs extend into
323 the range of “high CaO” MIs (≥ 13 wt.%; Schiano et al., 2000), which have been attributed to
324 localized dissolution-reaction and mixing processes (Danyushevsky et al., 2004). The high CaO
325 MIs could have formed from plagioclase dissolution, as they possess a positive Eu^* anomaly that
326 correlates with Al_2O_3 concentration (Fig. 4c). Importantly, the four MIs with the highest CO_2
327 derive from > 6 km depth and are part of the low SiO_2 MI population. These primitive compositions
328 are not observed in either the 1997 scoria (Zimmer et al. 2010) or the main July 12 phase of OK08
329 (Larsen et al., 2013), and we propose they originate from a unique reservoir (R_d). Given that some
330 of these MIs are hosted in unzoned crystals, R_d was possibly the source of the dike injection that
331 triggered the eruption. Given the range of Fo contents that host such low SiO_2 melts (Fo_{79-86}), the
332 range of saturation pressures (2–15 km, Fig. 8b) and diffusion timescales up to six months (e.g.,
333 RH-45, Fig 6b), R_d could reflect polybaric storage of inter-connected lenses of magma. Despite
334 being a unique reservoir, R_d shares the same parentage to the rest of Okmok, as all MIs and MG
335 and WR have overlapping Th/Nb (Fig. 4e).

336 ***6.2 Precursory seismicity related to magma ascent***

337 Previous studies emphasize Okmok’s shallow magma system, with seismic imaging
338 pointing to depths < 10 km and geodetic inflation depths <3 km. Our new homogenized MI CO_2
339 measurements point to an important, deeper mid-crustal magma storage region (R_d). Firstly, the
340 two MIs with the highest CO_2 content (RH22 and RH31-2, Fig. 7b and 7c) have saturation
341 pressures of 360 and 409 MPa, which translate to ~13 and ~15 km in depth, respectively. These

342 depths align closely with the deepest earthquakes in the five-hour precursory seismic swarm, ~15
343 km (Fig. 8c; Garza-Girón et al., 2023). Secondly, both crystals are unzoned in Fo (Fig. 7a and 7d)
344 and especially for RH-22 (Fo₈₆), which is out of equilibrium with the rims of all zoned olivine
345 (Fo₇₇₋₈₂), must have resided in the OK08 magma for a maximum of one day (Fig. 7e) – possible
346 within the five-hour seismic swarm. Lastly, RH-22 and RH-31 belong to the low SiO₂ MI
347 population unique to the initial ash (Fig. 3a). The appearance of the deep, primitive MI at the start
348 of the eruption supports a diking event to trigger OK08. At the same time, the shallower reservoirs
349 R_m and R_s were also tapped, evident from melt (the rim glass surrounding olivines from R_m and
350 the free-standing MG from R_s) and olivines (with low-CO₂ and bubble-free MIs from R_s). This
351 reinforces the case for a sudden run-up to OK08 and destabilization of several magma reservoirs.

352 Our results also indicate for the first time the presence of a CO₂-rich fluid at mid-crustal
353 depths beneath Okmok, which may facilitate dike initiation (via overpressure) and propagation
354 (via a gas-rich dike tip; Maimon et al., 2012), thus providing a mechanism for rapid vertical
355 transport to eruption. A magma ascent rate of 15 km in 5 hours, or 0.8 m/s, is typical of other
356 explosive eruptions (Cassidy et al., 2019). The point of CO₂-saturation has also been implicated
357 as a magma “launching depth,” based on the appearance of high-CO₂/S gas weeks before eruption
358 of open-vent volcanoes (Ding et al., 2025). Additionally, the distinct shift in the location of
359 hypocenters from the inter-eruptive to the pre-eruptive period (Power et al., 2019, Garza-Girón et
360 al., 2023; Fig. 1a) also suggests the new arrival of a new mid-crustal (15 km) stress source distinct
361 from the slow and steady source of shallow (3–4 km) geodetic inflation.

362 ***6.3 Linking petrological and geophysical events at Okmok***

363 Olivines with both reverse and normal zoning record timescales from weeks to years (using
364 DC07 diffusivities), indicating long-term injection of R_d olivines into into R_m or within R_d (normal

365 zoning) and movement of magma from R_m to R_s (reverse zoning). However, longer timescales are
366 predominantly recorded in reversely zoned crystals, while normally zoned crystals record shorter
367 timescales (Fig. 9a). This inference from the diffusion clocks agrees with the distribution obtained
368 by imaging the randomly selected olivine population, with P_1 (sharp) consisting of more normally
369 zoned crystals and P_3 (long) including more reversely zoned crystals (Fig. 5b). Importantly, there
370 is a ramp up in the number of zoned olivines towards the eruption. With reference to a population
371 of 52 timescales, an average of 0–2 olivines per month (<4% of the timescales) recording mixing
372 events between 2003 and 2007. This rate increases to 2–3 per month (<6%) in early 2008 and then
373 exponentially increases to 5 and 16 per month (~10% and 30%) in the two months before eruption.
374 These timescales provide important context to the evolution towards OK08.

375 The persistent appearance of reversely zoned crystals supports the hydraulic
376 interconnection of R_m and R_s hypothesized by Walwer et al. (2019) using deformation data and
377 resolved by Miller et al. (2020) and Ohlendorf et al. (2014) using seismic tomography. Apparently,
378 only the liquid from R_m is mixed into R_s as there is a lack of zoned crystals with Fo_{72} rims. The
379 longest reversely-zoned crystal, with a minimum timescale of 3.5 years, indicates that interactions
380 between R_m and R_s were happening well before the shift in the Mogi source 3-4 months prior to
381 eruption (Larsen et al., 2013). The lack olivine timescales coinciding with inflation events prior to
382 2005 could be related to the loss of resolution of the diffusion clock (Fig. 7b). In further support
383 of Walwer et al. (2019)'s hypothesis that oscillatory deformation is related to magma movement
384 between reservoirs, a portion of reversely zoned crystal timescales (20% $n=10$) falls within the
385 post-2006 period of oscillatory deformation, consistent with magma from R_m interacting with
386 olivines in R_s (Fig. 9a and 9b).

387 The increase in early 2008 of the number of mixing timescales per month coincides with
388 the initiation of the last inflation event (Fig. 9e). The subsequent ramp-up in the last two months
389 also coincides with the timing of predicted tensile failure of R_m (Albright et al., 2019) and shift in
390 the Mogi source (estimated to occur in early May; Freymueller and Kaufman (2010; Fig. 9f). The
391 destabilization of R_m could promote more intense mixing, as recorded increase in the number of
392 olivine timescales in the month prior to eruption.

393 Thus, there are several coincident events recorded in the olivine timescales (using DC07
394 diffusivities) and sensed geophysically, in the years, months and weeks prior to eruption of OK08.
395 If we instead use the $\sim 10x$ faster diffusivities from Shea et al. (2023), all the calculated olivine
396 timescales occur 7 months prior to OK08 (SI-17). 45 out of 52 timescales recorded in 43 olivines
397 would indicate magma mixing within 2 months of the eruption. Such results would be consistent
398 with the over-all ramp in events months prior to OK08 (Fig. 9). Use of the Shea et al. diffusivities
399 provide even less resolution in the olivine clocks to events in the years prior to OK08.

400

401 ***6.4 Temporal sequence of events leading up to the eruption***

402 Here we summarize the events preceding OK08, based on combined geochemical and
403 geophysical evidence. During the inter-eruptive period (1997–2008, Fig. 10), Okmok experienced
404 three to four inflation episodes (Fig. 9b) modeled as a single Mogi source at 3.4 km (Lu et al.,
405 2010; Wang et al., 2021); none of these inflation events were accompanied by heightened
406 seismicity (Bennington et al., 2015; Fig. 9d). Aseismic inflation has been attributed in other
407 systems to well-established melt channels feeding the main reservoir (Chouet & Matoza, 2013;
408 White et al., 2011). Concurrently, R_m actively mixes with R_s , evident from reversely zoned
409 crystals, barometry, seismic tomography and oscillatory deformation. R_d also shows mixing

410 amongst its polybaric storage regions, with one crystal, RH-45 (Fig 6b), containing primitive MI
411 compositions and reflecting magma mixing six months prior to OK08 at ~ 4 km.

412 In early 2008 (Fig. 10), the renewed inflation and shift in Mogi source coincides with a
413 rapid increase in the number of olivines recording mixing, suggesting more intense magma
414 movement between reservoirs. The possible overlap in timing of tensile failure and shift in the
415 Mogi source, together with the increased number of mixing events could indicate that the system
416 reached a critical state due to an increase in magma supply to a system already “primed” after 11
417 years of slow and steady inflation. However, we note that in modeling the occurrence of tensile
418 failure during that period, Albright et al. (2019) only consider R_m (and not the connection to R_s)
419 and do not consider the shift in the source identified by Freymueller and Kaufman (2010).

420 In the five-hours preceding the opening phase of OK08 (Fig. 10), the geophysical behavior
421 beneath Okmok changed to elevated seismicity. Inter-eruptive seismicity has been hypothesized
422 to be caused by reactivation of faults accommodating deformation (Johnson et al., 2010; Fig. 1a),
423 while the pre-eruptive seismicity has been linked to a planar feature opening in the mid-crust
424 (Garza-Girón et al., 2023; Fig. 10), which we have identified to be magmatic (R_d). The rapid
425 formation of the planar structure during the five-hour precursory seismicity links R_d to R_s and R_m ,
426 bringing along unzoned crystals from each reservoir (F_{072} , F_{077-80} , F_{086}). This small volume of
427 volatile rich magma from R_d may have destabilized the OK08 system and initiated the eruption.
428 While it is also possible that tensile failure of R_m led to depressurization that promoted magma
429 influx from depth, this hypothesis is difficult to test without any deformation data recorded in the
430 months preceding the eruption. There was also no obvious migration of the precursory-seismicity
431 with depth (Garza-Giron et al., 2023).

432 **7. Implications for Eruption Run-Up**

433 Here we have shown that the extremely short run-up for OK08 (5 hours) coincided with
434 rapid ascent of magma from depths comparable to the earthquakes (up to 15 km; Fig. 8a-c).
435 Moreover, the first magma erupted in OK08 was more mafic than the main eruptive phase that
436 followed for 4 hours (Fig. 3a), and contained olivine crystals that host MIs more CO-rich (Fig. 3b)
437 and primitive (Cr-rich; Fig. 4) than any other melts erupted from Okmok caldera over the past
438 2000 years. The mid-crustal source of crystals implicates a deeper magma storage region beneath
439 Okmok than had been recorded by prior geophysical studies (Fig. 8d). The inferred dike that
440 generated both the precursory seismicity and triggered the eruption forged a new vent in a location
441 8 km northeast of the vent that had hosted the prior century's eruptions (Fig. 1a). Thus, these
442 observations support a view that the short run-up was a new seismic event that was driven by
443 ascent of a new magma from mid-crustal depths in a new location than prior events, generating
444 the largest eruption at Okmok in over 2000 years. The short run-up of OK08 demonstrates how
445 sudden events may trigger eruptions and motivates the necessity for real-time seismic monitoring
446 that has been embraced by the Alaska Volcano Observatory (Power et al., 2020).

447 On the other hand, Okmok displayed over a decade of deformation prior to eruption (Fig.
448 1c), with episodic inflation of shallow magmatic reservoirs (≤ 3 km depth; Fig. 8d) that we have
449 shown were also tapped during OK08 (Figs. 8a, 10). Increasing uplift and approach to tensile
450 failure constitute a different kind of run-up, and one that we also relate to magma input (Fig. 9a-
451 c). With olivine zonation timescales, we document an increase input of magma from deep to
452 shallow storage regions in the months before eruption (Fig. 9e). Apparently, Okmok was already
453 moving toward failure (e.g., it was "primed") when the sudden input of deep, volatile-rich magma
454 triggered eruption. This scenario provides a more sanguine view for forecasting. The success of

455 the Albright et al. (2019) hindcast model for OK08, based entirely on InSAR and GPS data, holds
456 promise for providing warning when inflating systems like Okmok approach failure.

457 Roman and Cashman (2018) present several examples of “top-down” seismic behavior
458 observed at arc volcanoes, whereby shallow precursory seismicity gives way to deeper seismicity
459 that extends to the main magma storage region once the eruption is underway. OK08 represents a
460 counter example of a volcano with “bottom-up” behavior, whereby seismicity deeper than the main
461 magma storage region occurs prior to the eruption. The different depths (Fig. 8a) and short
462 residence time of crystals (Fig. 9e) that we have documented provide compelling evidence that
463 OK08’s “bottom-up” precursor was related to the input of deep mafic magma. The causes of “top-
464 down” vs “bottom-up” behavior are as yet unclear. Our study highlights the value of
465 retrospectively integrating petrology with monitoring data to identify magmatic processes
466 underlying precursory signals.

467 Ours is one of a growing number of studies demonstrating that arc magmas are richer in
468 CO₂ than previously thought (>1000s ppm; Ding et al., 2025; Rasmussen et al., 2020; Taracsák et
469 al., 2023). CO₂ exsolution has been proposed to drive high decompression rates, resulting in
470 explosive eruptions (Allison et al., 2021; Barth et al., 2024; DeVitre et al., 2023). The first four
471 hours of OK08 erupted an estimated 0.048 km³ volume of magma (Unema et al., 2016),
472 corresponding to an approximate mass eruption rate of ~10⁷ kg/s and a decompression rate of 10
473 m/s – similar to those at Fuego 1974 and Etna Fall Stratified, other VEI 4 mafic eruptions (Barth
474 et al., 2019, 2024). Thus, OK08 is consistent with the emerging view that decompression rate, and
475 not water concentration, scales with the intensity and magnitude of mafic eruptions (Ferguson et
476 al., 2016; Moussallam et al., 2019). The rapid ascent of magma and the high CO₂ in OK08 MIs
477 (>1800 ppm) suggest that the “launching” depth (Ding et al., 2025), where CO₂ exsolves, may also

478 be an important for eruption triggering in closed-system volcanoes such as Okmok. This premise
479 calls for a re-examination of decades of MI studies that did not account for CO₂ in the vapor
480 bubble. Future studies may explore the physical relationship between the CO₂ content of magmas
481 and the speed of dike propagation; such a relationship could directly link geochemical, geophysical
482 and eruptive processes that lead to explosive eruptions with short run-up.

483

484 **Figure Captions**

485 Figure 1. a) Map of Okmok caldera using the Digital Elevation Model by Schaefer et al. (2012)
486 and earthquake data from Power et al. (2019). Inter-eruptive seismicity represents events before
487 12 July 2008 and pre-eruptive seismicity represents events on 12 July 2008 before the eruption.
488 Dashed line represents cross section in Figure 10, and is drawn through Cone A (site of 20th century
489 eruptions), Ahmanalix (new cone formed in 2008, near older Cone D). b) A map of Alaska and
490 the Aleutian arc, the red box indicate the location of Okmok. c) Timeseries of volume change of
491 Okmok's magma body since the availability of InSAR data in 1992. Data extracted from Lu et al.
492 (2014), Fournier et al. (2009), Wang et al. (2021), Xue et al. (2020). Post 2008 data for Lu et al.
493 (2014) extracted from National Academies of Sciences, Engineering, and Medicine (2017) report.

494

495 Figure 2. a) FTIR spectra of RH22, one of the high CO₂ MIs. Baselines obtained from PyIRoGlass
496 (Shi et al., 2024). b) Transmission microscope image of bubble-bearing MI before and after
497 experimental rehomogenization. Top row is a transmission microscope image of a melt inclusion,
498 MI1, hosted in RH43 and bottom row is a transmission microscope image of a melt inclusion, MI2,
499 hosted in RH31. c) Transmission microscope image of an olivine, RH23, with a doubly-polished
500 (MI1) and two other (MI2 & MI3) rehomogenized MIs.

501

502 Figure 3. a) PEC-corrected melt inclusion (MI) and matrix glass (MG) compositions measured in
503 this study from the initial phase of OK08. Untreated MIs that are bubble-free have a black border
504 while bubble-bearing have no border. MG (olivine) represent matrix glass attached to an olivine
505 crystal while MG (free) represent matrix glass picked directly from the ash. Fields are literature
506 data from Larsen et al. (2013). b) H₂O vs CO₂ with isobars calculated from MagmaSat (using
507 composition of RH22, which has the highest CO₂). Samples plotted in the yellow region are below
508 the detection limit of the FTIR. See text and Supplementary Table S1 for detection limits (in all
509 cases < 100 ppm).

510

511 Figure 4. Comparison of major and trace element compositions of OK08 melt inclusions and
512 matrix glass from this study to literature whole rock data. Whole rock data from GEOROC, Finney
513 et al., (2008), Larsen et al., (2013). Black solid line represent inferred LLD a) Cr against SiO₂
514 with labels for the MIs' olivine host Fo content. There is a single data point (arrow) for RH-22 at
515 394 ppm Cr. b) V against SiO₂. Mixing lines is for the average cluster of MG (Free) and the most
516 primitive homogenized MI. Arrows indicating the inferred liquid line of descent and 'kink' after
517 magnetite starts to crystallize. c) Eu/Eu* against Al₂O₃. Positive Eu anomaly where Eu/Eu* > 1
518 for whole rocks suggest plagioclase accumulation while for melt inclusions suggest plagioclase
519 dissolution. Negative Eu anomaly where Eu/Eu* < 1 for whole rock suggest plagioclase removal
520 while for melt inclusions suggest plagioclase crystallization. d) Sc against SiO₂ e) Th against Nb

521

522

523 Figure 5. a) Histogram of olivine Fo compositions for initial ash of OK08. Symbols on top
524 represent the glass and whole rock equilibrium Fo content, assuming $Fe^{3+}/Fe^T = 0.2$ and $K_D = 0.3$
525 (symbols the same as Figure 3 and 4). The 1958, 1997, 2008 main phase and 2008 initial ash whole
526 rock data, purple, white, yellow and blue stars respectively, are obtained from this study, Finney
527 et al. (2008), Larsen et al. (2013) and Zimmer et al. (2010). b) Populations derived from backscatter
528 electron profiles for 72 different olivines from 300 – 1mm in size. See text for the description of
529 the different populations. c) Ponytail plot of the zonation profiles modelled for diffusion timescales
530 in this study. The colors represents the timescales modelled using the expression for Fe-Mg
531 diffusion in olivine from Dohmen and Chakraborty (2008).

532

533 Figure 6. Fo and P_2O_5 profiles of three representative olivines (a: DF05, b: RH45, c: DF07) with
534 increasing diffusion timescales from a) weeks to b) months to c) years. BSE images on the right
535 show orientations and profiles (red lines). The similar lengthscale for a) and b) olivines but an
536 order of magnitude difference in time is due to the orientation of the profile. a) is almost parallel
537 to the c-axis [001] which has diffusivities 6x faster than the other two axes. Red solid lines
538 represent fit to homogenous initial conditions while blue solid lines represent fit to an initial step
539 profile. Initial conditions are shown as dashed lines, for red and blue models, respectively. t_{DC}
540 represents timescale using diffusivities by Dohmen and Chakraborty (2008) while t_s represents
541 timescale using diffusivities 10x higher as estimated by Shea et al. (2023).

542

543 Figure 7. Images, profiles and models of olivines with the highest CO₂ MIs. a) Backscatter electron
544 image of RH31. Transmitted image of b) RH22 and c) RH31. d) Fo profiles of RH22 and RH31
545 as indicated with colored lines in b) and c). e) Results from a forward 1-D finite difference
546 diffusion model, with profiles along [100] dashed and [100] solid, in order to estimate the time to
547 observe no Fo zoning on the EPMA (assuming no resolution ≤ 5 microns from the rim). f) Results
548 from a forward 3-D finite difference model of crystal sizes 500 μm (dashed) and 1 mm (solid).
549 The 1-D profile is obtained using a [100]-[001] slice through the core of the crystal along [001]
550 axis. A 500 μm initially homogenous crystal at Fo₇₂ will equilibrate its core composition with Fo₈₀
551 in 5-11 years at T of 1115°C.

552

553 Figure 8. a) Melt inclusion and matrix glass data from this study for OK08 and Zimmer et al.
554 (2010) for the 1997 eruption. Symbols follow previous figures. The estimated depth of the evolved
555 reservoir is represented by the yellow star, using the SiO₂ concentration and equilibrium Fo of
556 the 2008 (main) whole rock. Arrow is drawn to guide the eye when comparing the range of depths
557 vs. composition of melt inclusions. a) SiO₂ of olivine hosted MIs and b) Fo content of their olivine
558 host plotted at their MIs vapor saturation depth. c) Earthquakes (EQs) binned by 1 km (Garza-
559 Giron et al. (2023) five-hours pre-eruptive seismicity). d) Compilation of storage regions identified
560 using seismic (red) and geodetic data (green). Data from Mann et al. (2002), Lu et al. (2010), Lu
561 and Dzurisin (2010), Xue et al. (2020), Miller et al. (2020), Kasatkina et al. (2022), Ohlendorf et
562 al. (2014). Description for Miller et al. (2020) and Kasatkina et al. (2022) are as given in those
563 respective studies.

564

565 Figure 9. A comparison between various temporal geophysical and geochemical data in the five
566 years prior to the 2008 eruption of Okmok. Blue region represents deflation, red represents
567 inflation and grey represents the shift in the Mogi source observed by Freymueller and Kaufmann
568 (2010) and during the time period with limited GPS data. a) Rank order plot of diffusion timescales
569 (symbols), using diffusivities in Dohmen and Chakraborty (2008). Colors are same as Figure 5. S
570 represent the diffusion timescales of inner shoulder profiles. Unzoned profiles are placed at time
571 0 with an error of 1 day. Normally zoned profiles are shown in filled symbols, reversely zoned
572 profiles are open. Histograms plot number of olivine timescales/month in the 52 sample
573 population. b) Deformation data from two GPS stations in the caldera (see Figure 1 for locations).
574 Scale is relative to the first data point. OKCD data from Nevada Geodetic Lab (NGL); OKCE from
575 Alaska Volcano Observatory. c) The mean tensile stress and 2σ from joint inversion in Albright et
576 al. (2019). The red line represents 0 MPa tensile strength which is the lowest threshold in Albright
577 et al. (2019). d) Earthquake data using the Power et al. (2019) catalogue. Inter-eruptive and pre-
578 eruptive classification follows Figure 1. e) and f) show data from a year before OK08. e) is a
579 zoomed-in version of a). Arrows shows the ramp up in the number of mixing events recorded in
580 olivines per month. f) is a zoomed-in version of b) and c) replotted together. Arrows indicate
581 notable geodetic events. The uncertainty of the timing of the shift in Mogi source is indicated by
582 the range of arrows (due to a lack of GPS data in May 2008).

583

584

585 Figure 10. A model of processes occurring beneath Okmok from the 1997 to the 2008 eruption.
586 The cross-section is SW-NE across the caldera (See dotted line in Fig 1.). Key geochemical and
587 geophysical observations for each time period are emphasized. SiO_2 content of the melt and Fo of
588 the olivine are scaled by intensity of red and green respectively to show the diverse compositions
589 in OK08, and where they are likely placed in the entire plumbing architecture. Stars represent
590 seismicity during the period.
591

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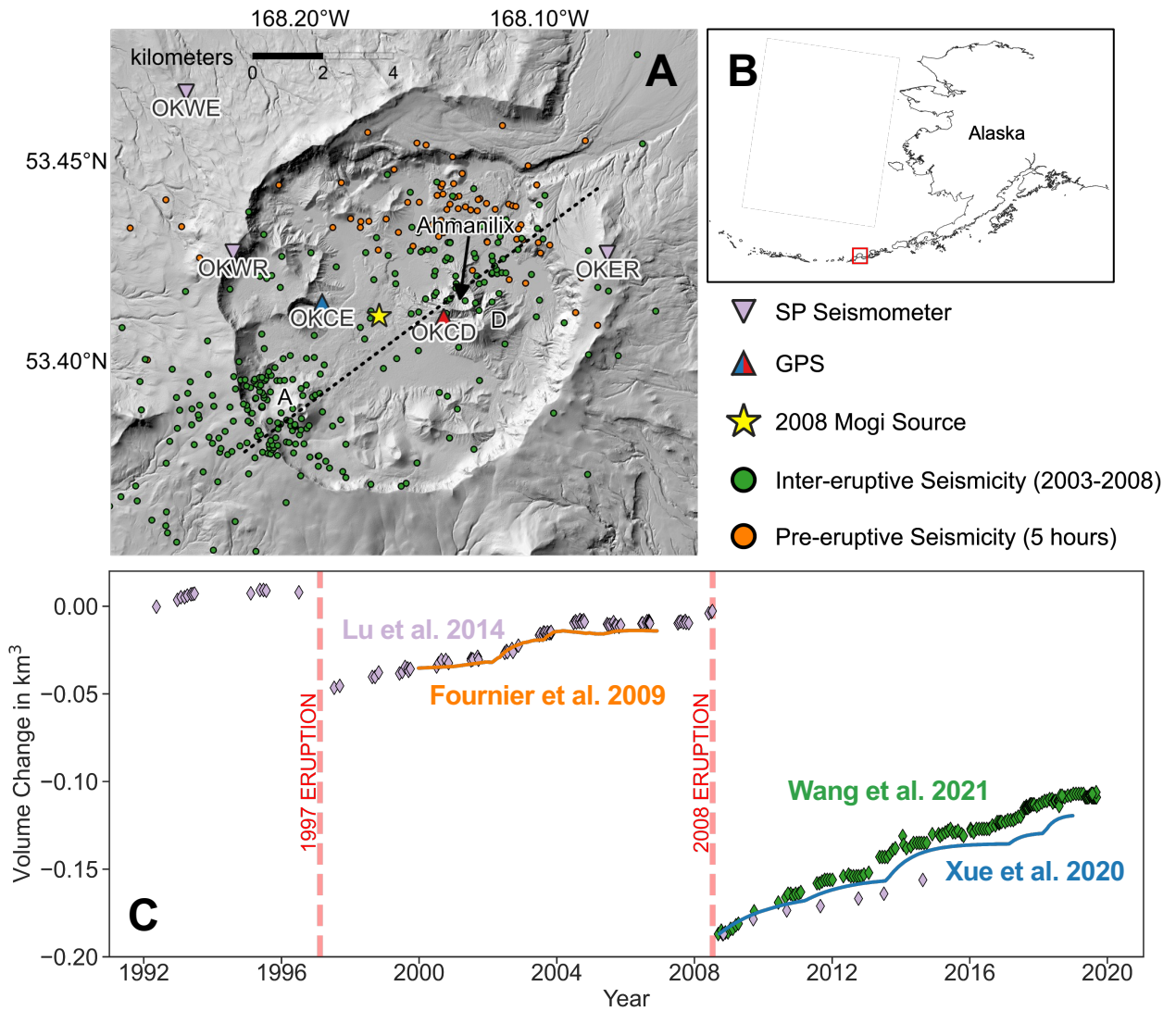


Figure 1. a) Map of Okmok caldera using the Digital Elevation Model by Schaefer et al. (2012) and earthquake data from Power et al. (2019). Inter-eruptive seismicity represents events before 12 July 2008 and pre-eruptive seismicity represents events on 12 July 2008 before the eruption. Dashed line represents cross section in Figure 10, and is drawn through Cone A (site of 20th century eruptions), Ahmanilix (new cone formed in 2008, near older Cone D). b) A map of Alaska and the Aleutian arc, the red box indicate the location of Okmok. c) Timeseries of volume change of Okmok's magma body since the availability of InSAR data in 1992. Data extracted from Lu et al. (2014), Fournier et al. (2009), Wang et al. (2021), Xue et al. (2020). Post 2008 data for Lu et al. (2014) extracted from National Academies of Sciences, Engineering, and Medicine (2017) report.

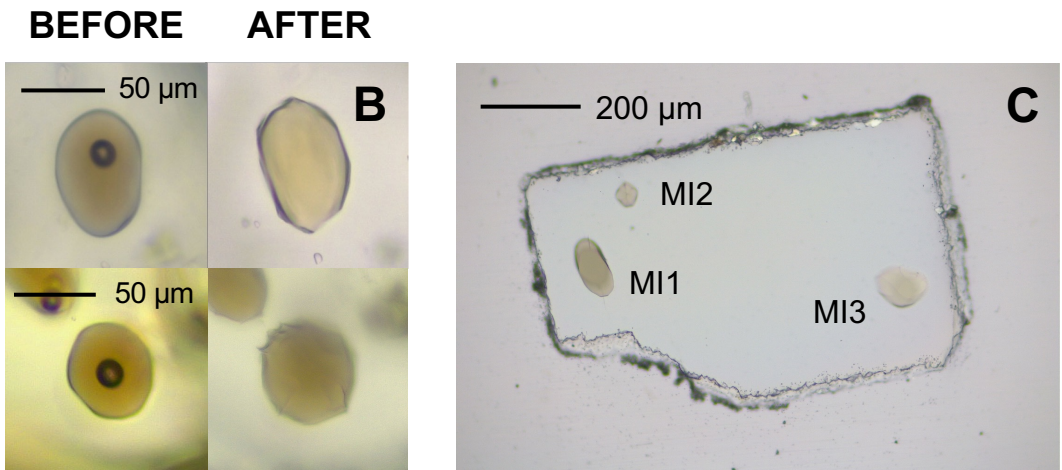
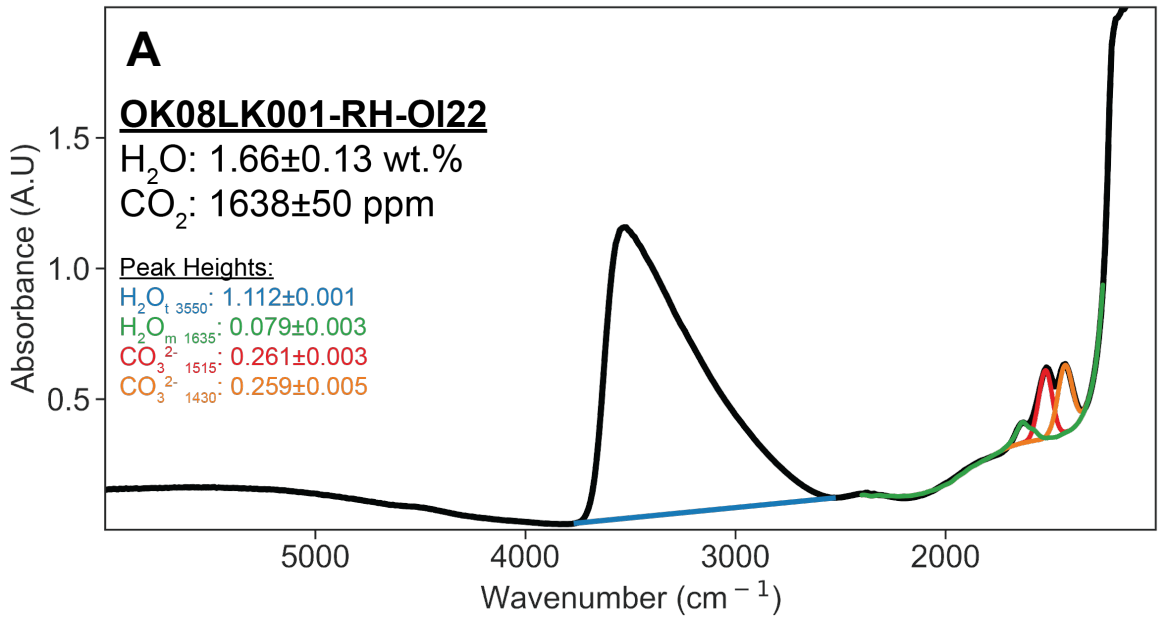


Figure 2. a) FTIR spectra of RH22, one of the high CO₂ MIs. Baselines obtained from PyIRoGlass (Shi et al., 2024). b) Transmission microscope image of bubble-bearing MI before and after experimental rehomogenization. Top row is a transmission microscope image of a melt inclusion, MI1, hosted in RH43 and bottom row is a transmission microscope image of a melt inclusion, MI2, hosted in RH31. c) Transmission microscope image of an olivine, RH23, with a doubly-polished (MI1) and two other (MI2 & MI3) rehomogenized MIs.

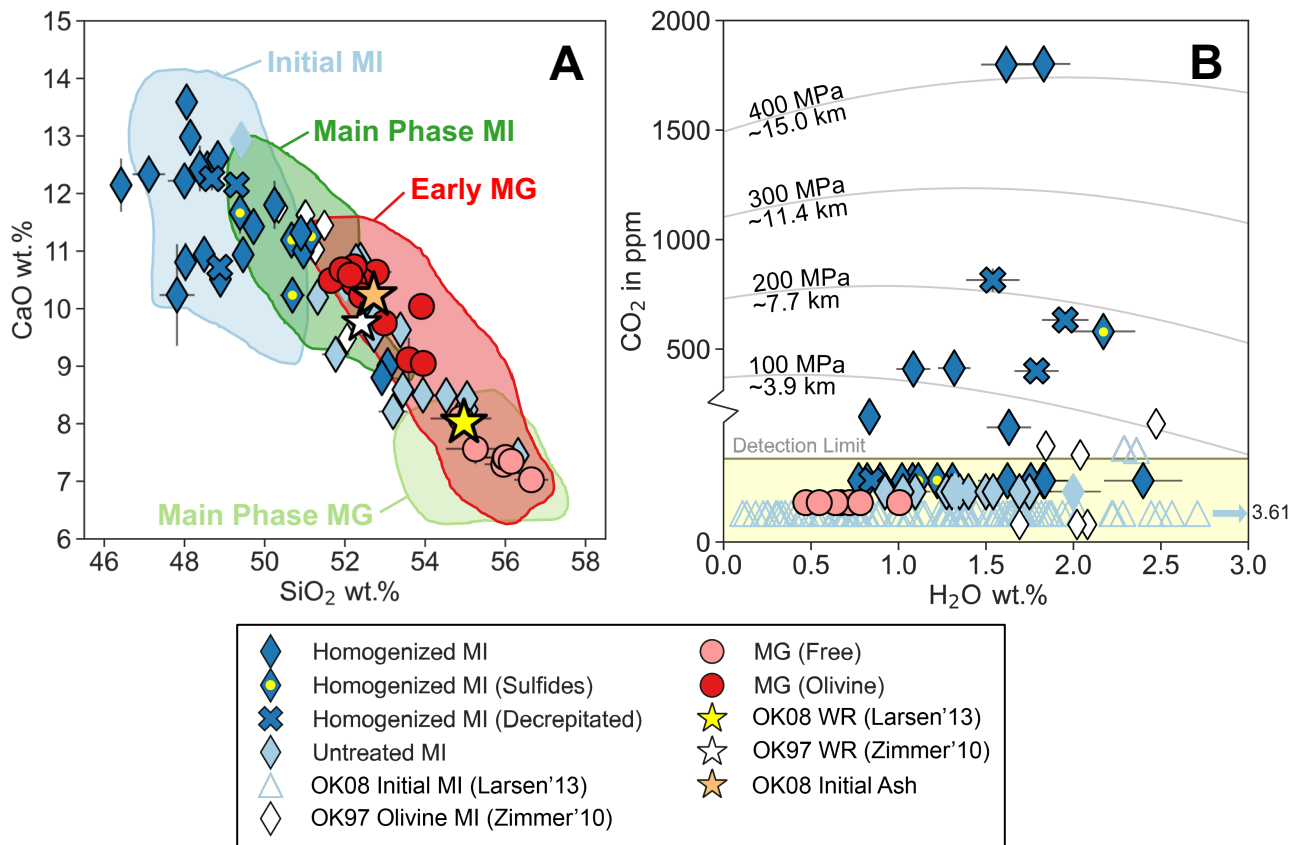


Figure 3. a) PEC-corrected melt inclusion (MI) and matrix glass (MG) compositions measured in this study from the initial phase of OK08. Untreated MIs that are bubble-free have a black border while bubble-bearing have no border. MG (olivine) represent matrix glass attached to an olivine crystal while MG (free) represent matrix glass picked directly from the ash. Fields are literature data from Larsen et al. (2013). b) H₂O vs CO₂ with isobars calculated from MagmaSat (using composition of RH22, which has the highest CO₂). Samples plotted in the yellow region are below the detection limit of the FTIR. See text and Supplementary Table S1 for detection limits (in all cases < 100 ppm).

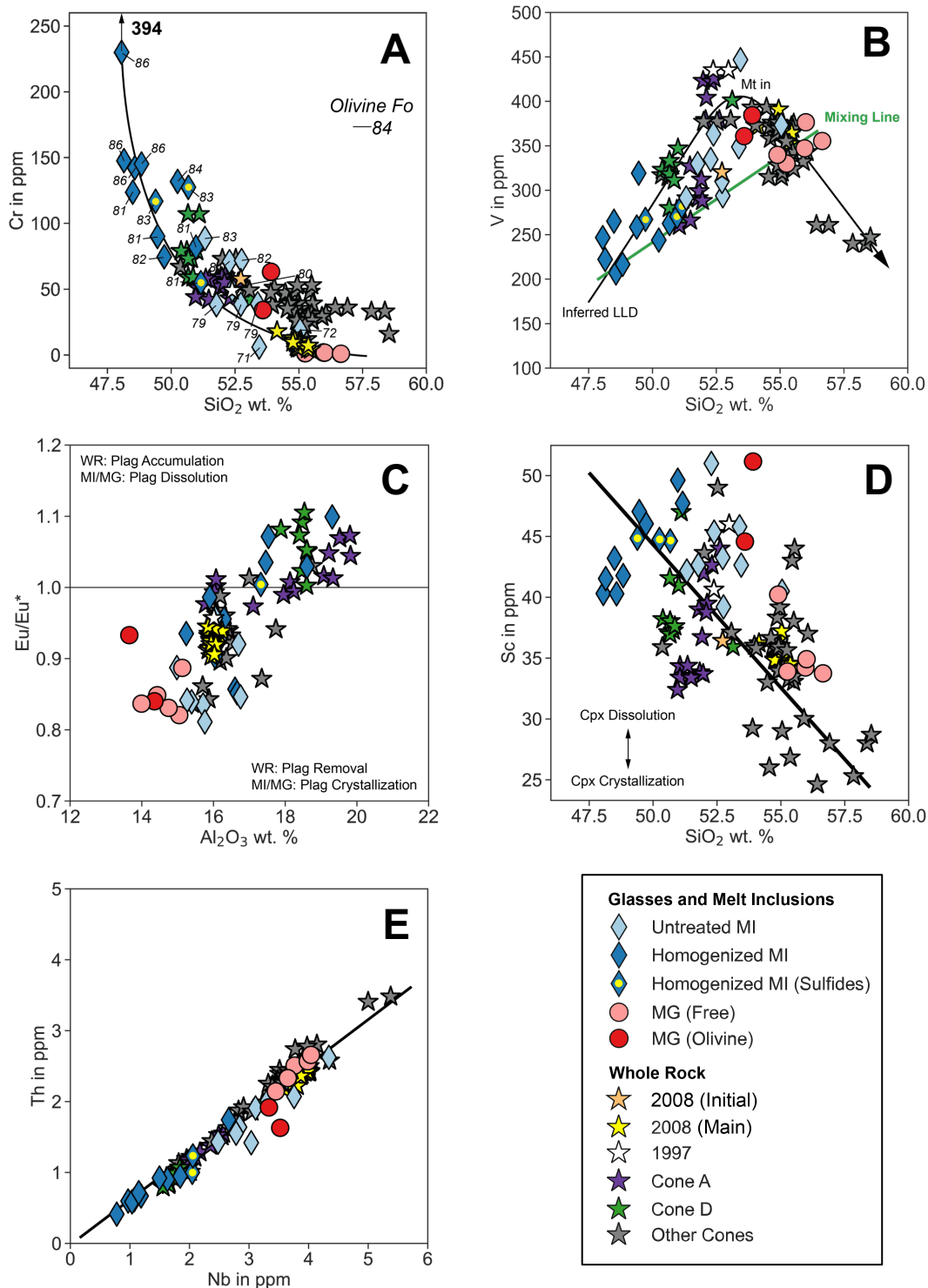


Figure 4. Comparison of major and trace element compositions of OK08 melt inclusions and matrix glass from this study to literature whole rock data. Whole rock data from GEOROC, Finney et al., (2008), Larsen et al., (2013). Black solid line represent inferred LLD a) Cr against SiO₂ with labels for the MIs' olivine host Fo content. There is a single data point (arrow) for RH-22 at 394 ppm Cr. b) V against SiO₂. Mixing lines is for the average cluster of MG (Free) and the most primitive homogenized MI. Arrows indicating the inferred liquid line of descent and 'kink' after magnetite starts to crystallize. c) Eu/Eu* against Al₂O₃. Positive Eu anomaly where Eu/Eu* > 1 for whole rocks suggest plagioclase accumulation while for melt inclusions suggest plagioclase dissolution. Negative Eu anomaly where Eu/Eu* < 1 for whole rock suggest plagioclase removal while for melt inclusions suggest plagioclase crystallization. d) Sc against SiO₂ e) Th against Nb

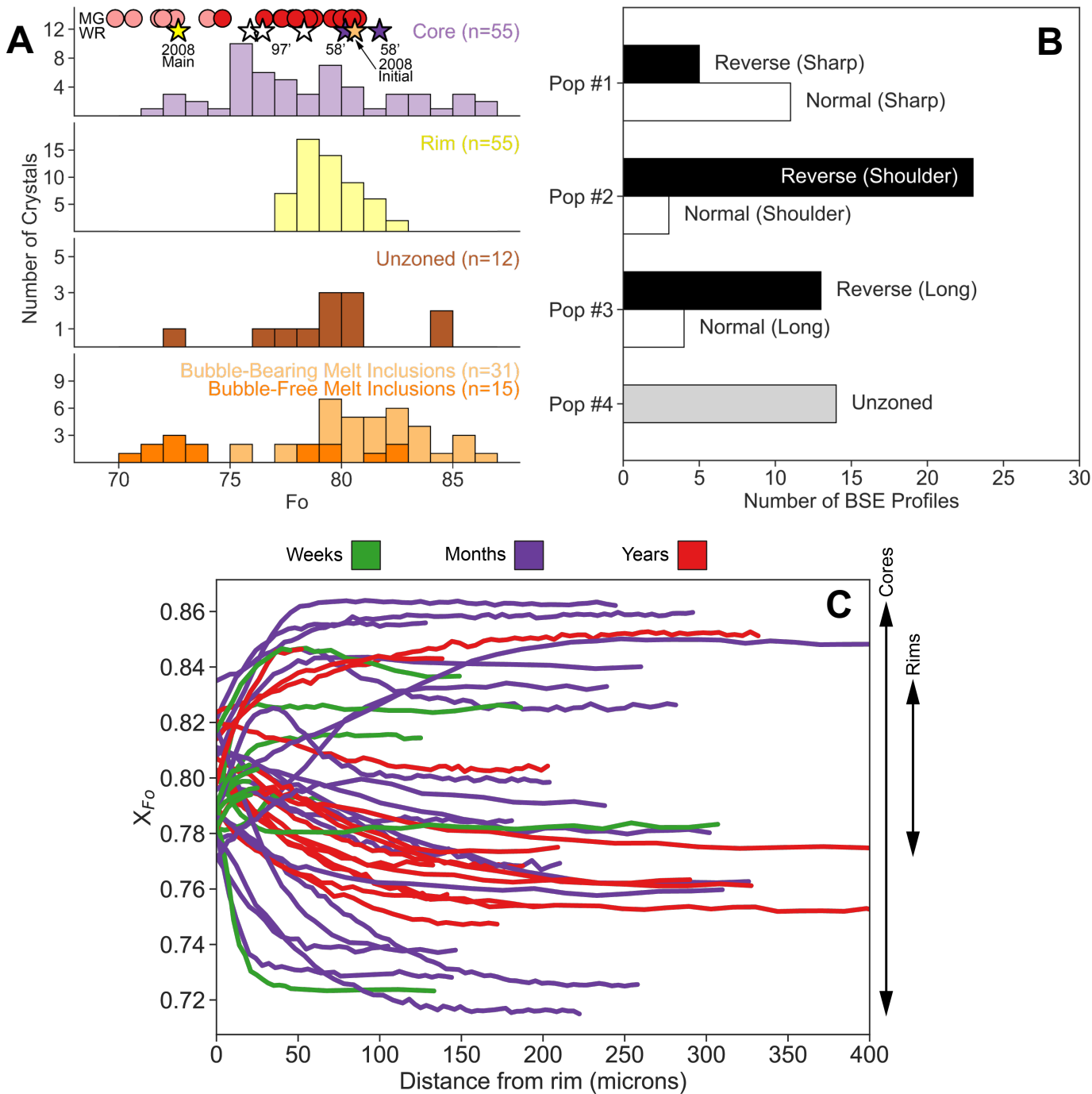


Figure 5. a) Histogram of olivine Fo compositions for initial ash of OK08. Symbols on top represent the glass and whole rock equilibrium Fo content, assuming $Fe^{3+}/Fe^T = 0.2$ and $K_D = 0.3$ (symbols the same as Figure 3 and 4). The 1958, 1997, 2008 main phase and 2008 initial ash whole rock data, purple, white, yellow and blue stars respectively, are obtained from this study, Finney et al. (2008), Larsen et al. (2013) and Zimmer et al. (2010). b) Populations derived from backscatter electron profiles for 72 different olivines from 300 – 1mm in size. See text for the description of the different populations. c) Ponytail plot of the zonation profiles modelled for diffusion timescales in this study. The colors represents the timescales modelled using the expression for Fe-Mg diffusion in olivine from Dohmen and Chakraborty (2008).

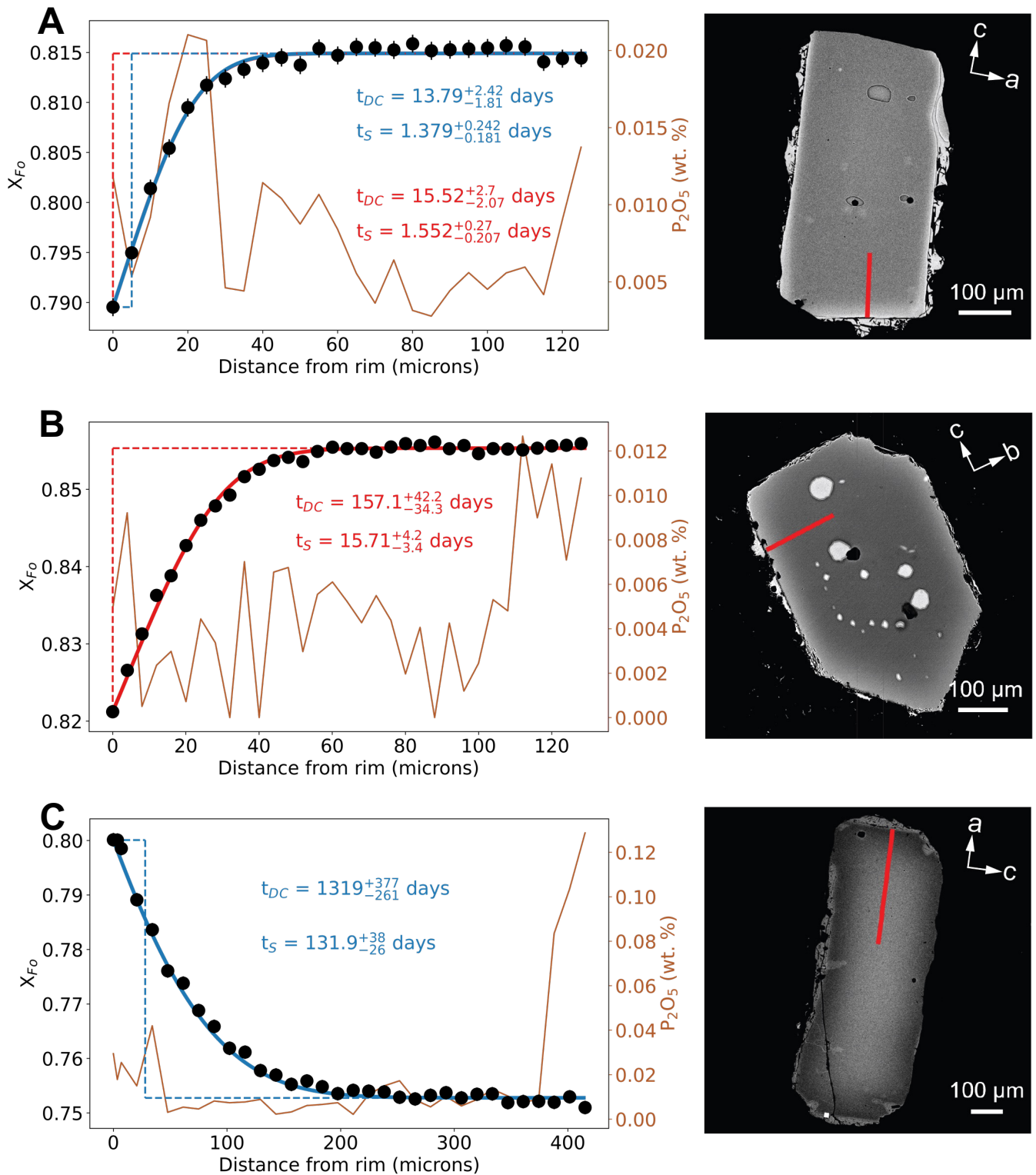


Figure 6. Fo and P_2O_5 profiles of three representative olivines (a: DF05, b: RH45, c: DF07) with increasing diffusion timescales from a) weeks to b) months to c) years. BSE images on the right show orientations and profiles (red lines). The similar lengthscale for a) and b) olivines but an order of magnitude difference in time is due to the orientation of the profile. a) is almost parallel to the c-axis [001] which has diffusivities 6x faster than the other two axes. Red solid lines represent fit to homogenous initial conditions while blue solid lines represent fit to an initial step profile. Initial conditions are shown as dashed lines, for red and blue models, respectively. t_{DC} represents timescale using diffusivities by Dohmen and Chakraborty (2008) while t_S represents timescale using diffusivities 10x higher as estimated by Shea et al. (2023).

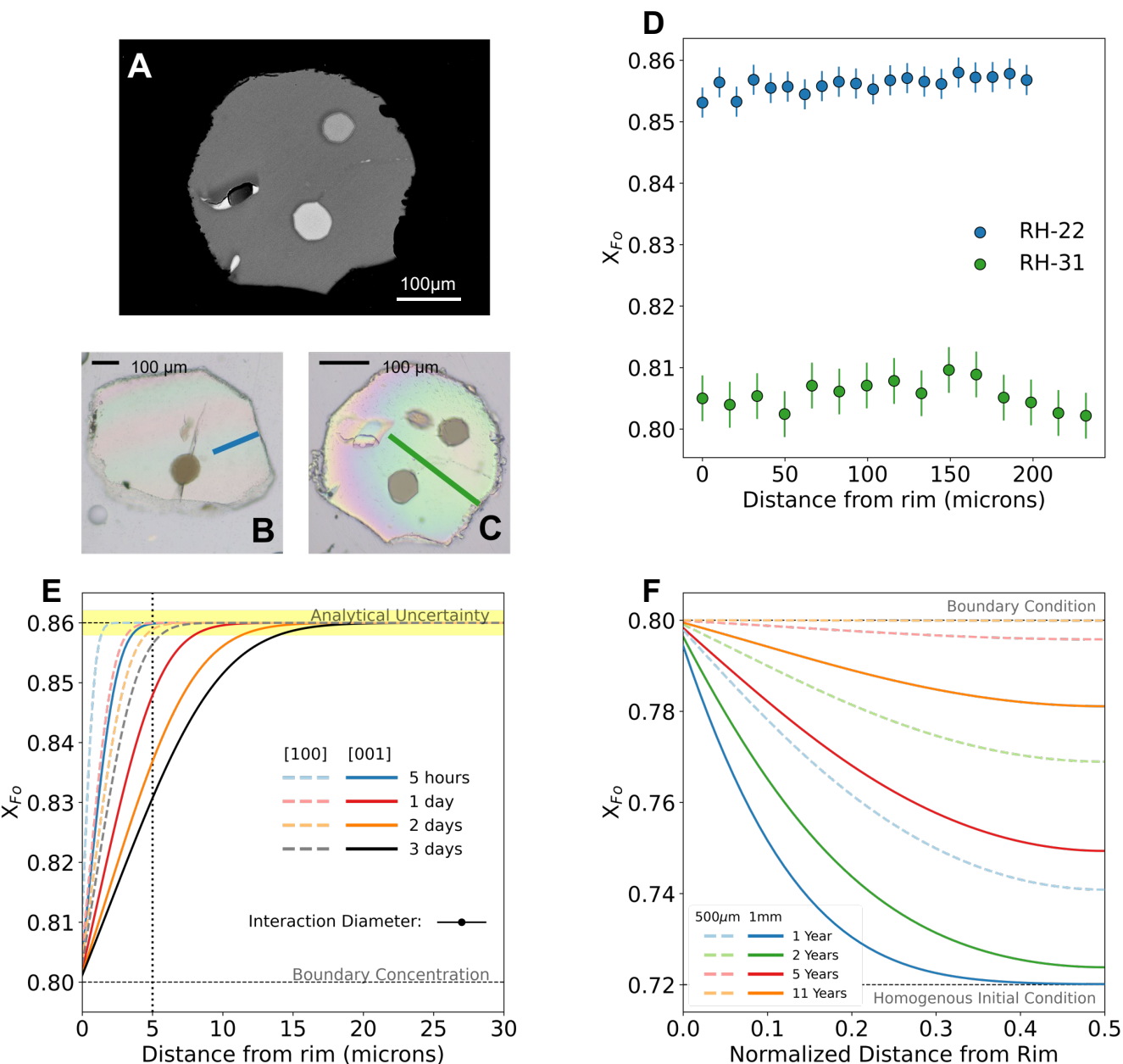


Figure 7. Images, profiles and models of olivines with the highest CO₂ MIs. a) Backscatter electron image of RH31. Transmitted image of b) RH22 and c) RH31. d) Fo profiles of RH22 and RH31 as indicated with colored lines in b) and c). e) Results from a forward 1-D finite difference diffusion model, with profiles along [100] dashed and [001] solid, in order to estimate the time to observe no Fo zoning on the EPMA (assuming no resolution ≤ 5 microns from the rim). f) Results from a forward 3-D finite difference model of crystal sizes 500 μm (dashed) and 1 mm (solid). The 1-D profile is obtained using a [100]-[001] slice through the core of the crystal along [001] axis. A 500 μm initially homogenous crystal at Fo₇₂ will equilibrate its core composition with Fo₈₀ in 5-11 years at T of 1115°C.

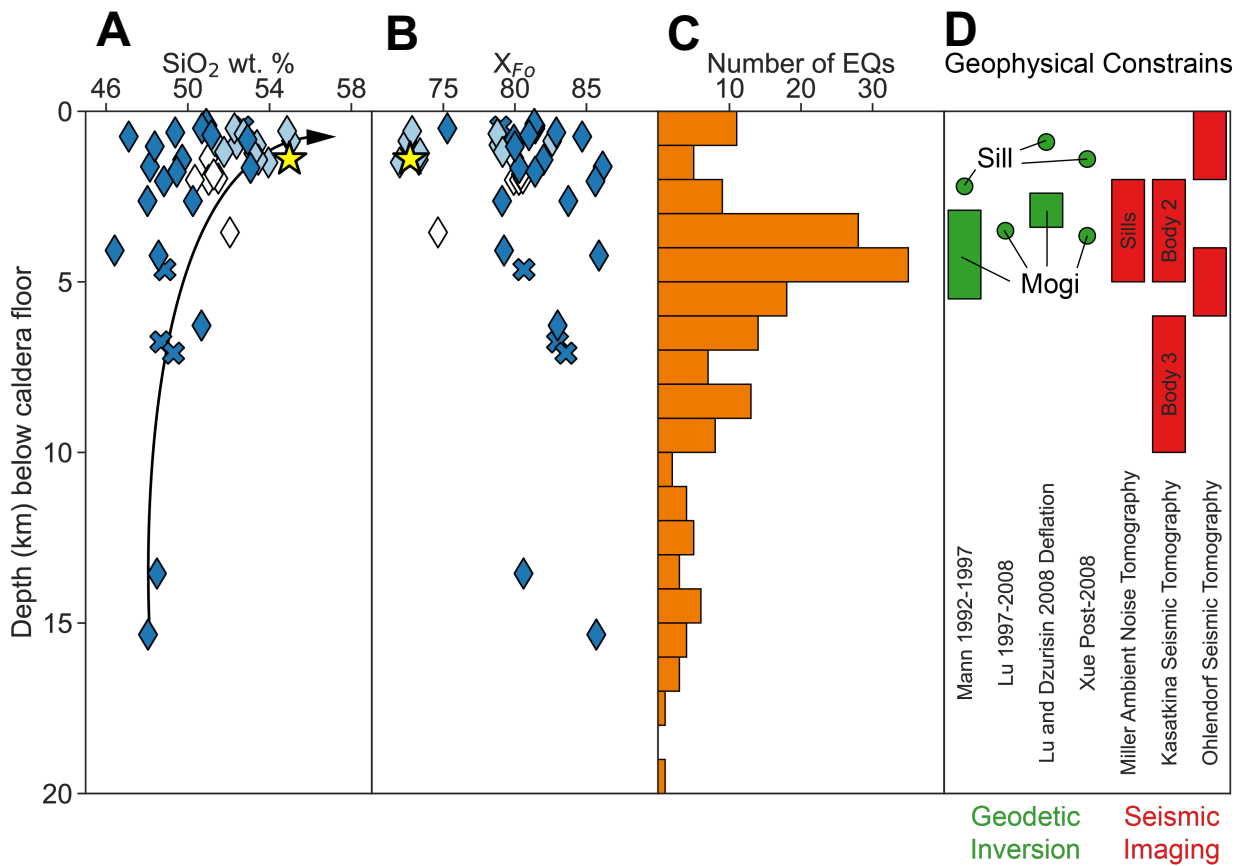


Figure 8. a) Melt inclusion and matrix glass data from this study for OK08 and Zimmer et al. (2010) for the 1997 eruption. Symbols follow previous figures. The estimated depth of the evolved reservoir is represented by the yellow star, using the SiO₂ concentration and equilibrium Fo of the 2008 (main) whole rock. Arrow is drawn to guide the eye when comparing the range of depths vs. composition of melt inclusions. a) SiO₂ of olivine hosted MIs and b) Fo content of their olivine host plotted at their MIs vapor saturation depth. c) Earthquakes (EQs) binned by 1 km (Garza-Giron et al. (2023) five-hours pre-eruptive seismicity). d) Compilation of storage regions identified using seismic (red) and geodetic data (green). Data from Mann et al. (2002), Lu et al. (2010), Lu and Dzurisin (2010), Xue et al. (2020), Miller et al. (2020), Kasatkina et al. (2022), Ohlendorf et al. (2014). Description for Miller et al. (2020) and Kasatkina et al. (2022) are as given in those respective studies.

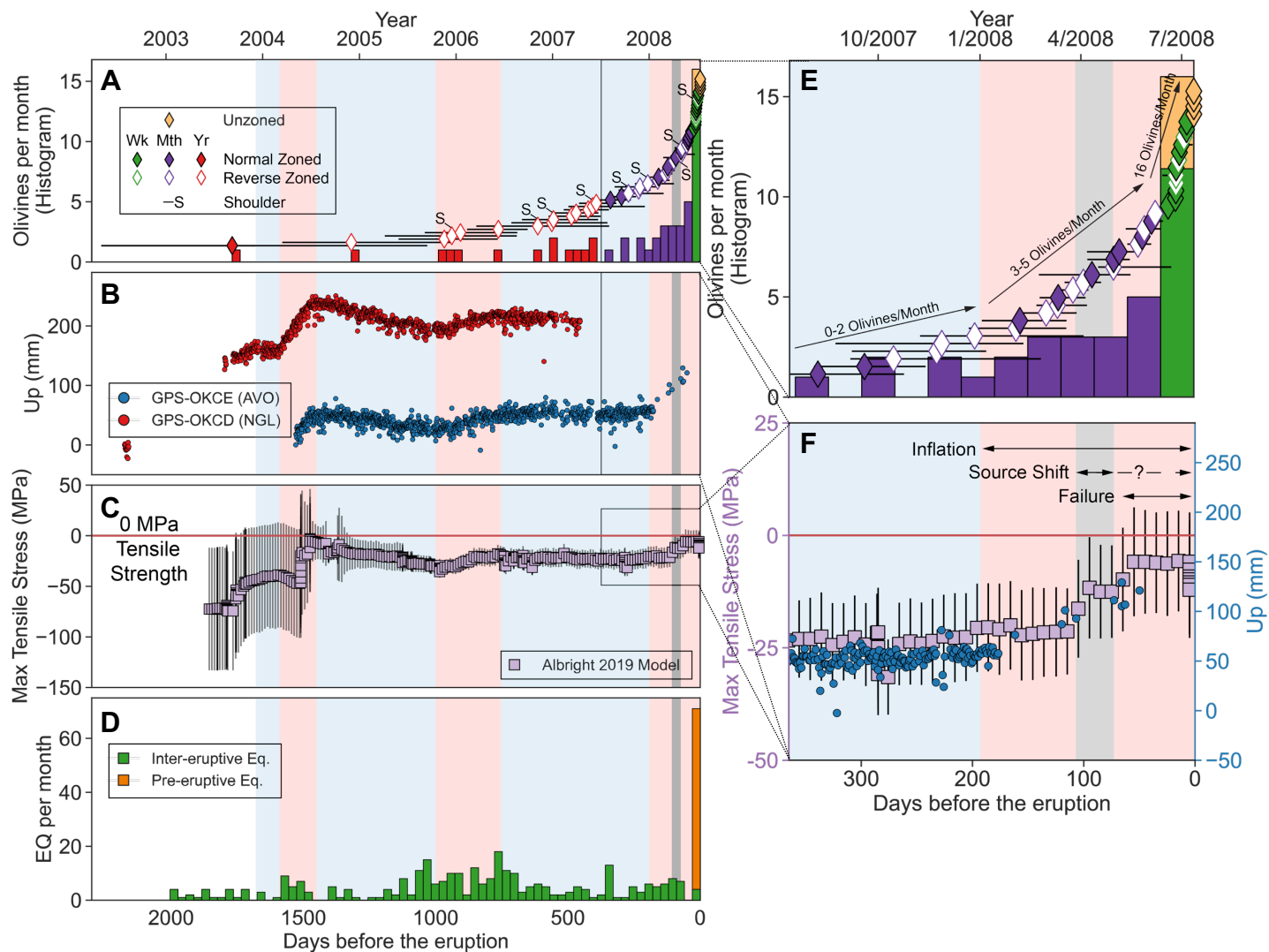


Figure 9. A comparison between various temporal geophysical and geochemical data in the five years prior to the 2008 eruption of Okmok. Blue region represents deflation, red represents inflation and grey represents the shift in the Mogi source observed by Freymueller and Kaufmann (2010) and during the time period with limited GPS data. a) Rank order plot of diffusion timescales (symbols), using diffusivities in Dohmen and Chakraborty (2008). Colors are same as Figure 5. S represent the diffusion timescales of inner shoulder profiles. Unzoned profiles are placed at time 0 with an error of 1 day. Normally zoned profiles are shown in filled symbols, reversely zoned profiles are open. Histograms plot number of olivine timescales/month in the 52 sample population. b) Deformation data from two GPS stations in the caldera (see Figure 1 for locations). Scale is relative to the first data point. OKCD data from Nevada Geodetic Lab (NGL); OKCE from Alaska Volcano Observatory. c) The mean tensile stress and 2σ from joint inversion in Albright et al. (2019). The red line represents 0 MPa tensile strength which is the lowest threshold in Albright et al. (2019). d) Earthquake data using the Power et al. (2019) catalogue. Inter-eruptive and pre-eruptive classification follows Figure 1. e) and f) show data from a year before OK08. e) is a zoomed-in version of a). Arrows shows the ramp up in the number of mixing events recorded in olivines per month. f) is a zoomed-in version of b) and c) replotted together. Arrows indicate notable geodetic events. The uncertainty of the timing of the shift in Mogi source is indicated by the range of arrows (due to a lack of GPS data in May 2008).

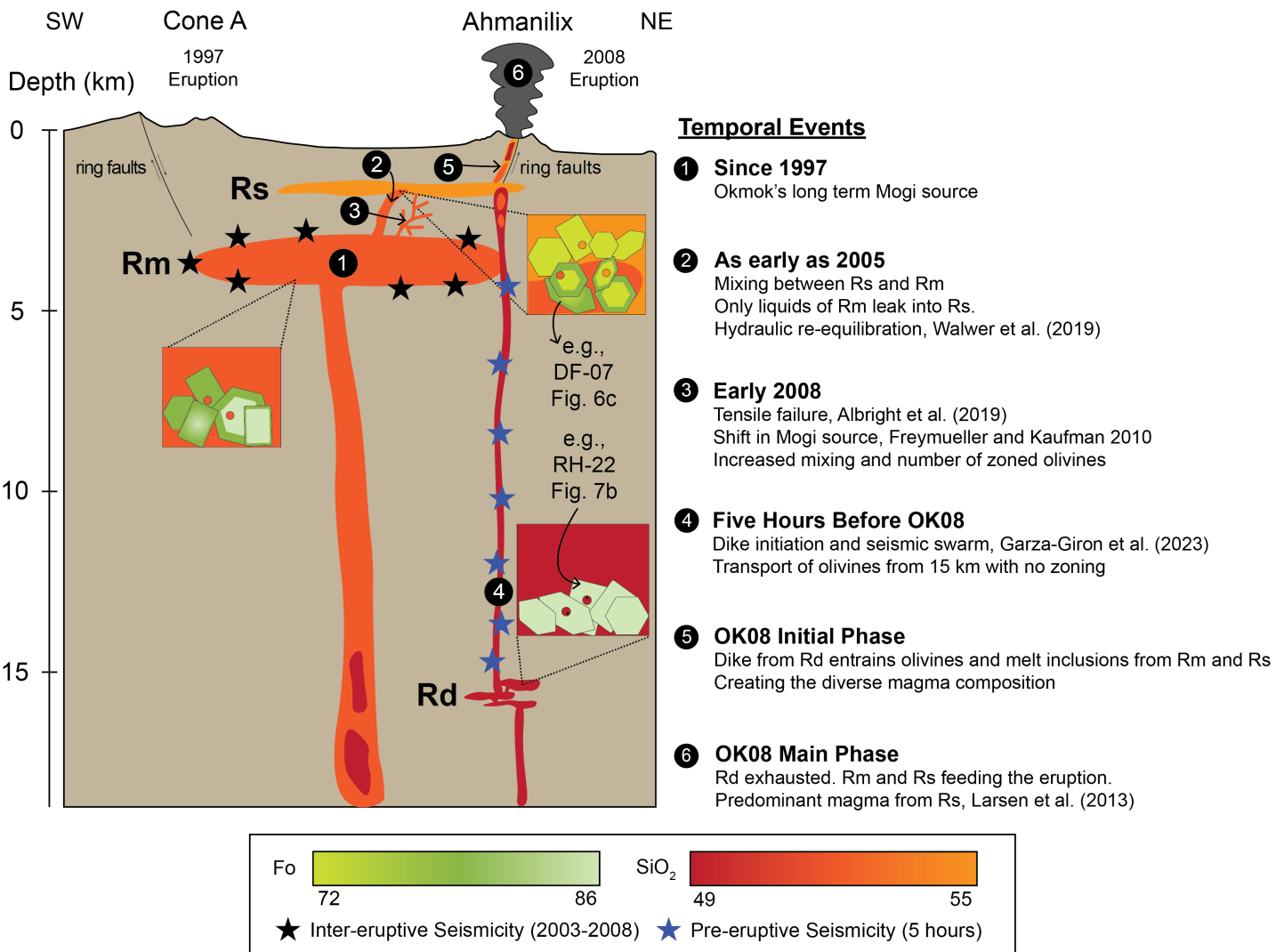


Figure 10. A model of processes occurring beneath Okmok from the 1997 to the 2008 eruption. The cross-section is SW-NE across the caldera (See dotted line in Fig 1.). Key geochemical and geophysical observations for each time period are emphasized. SiO₂ content of the melt and Fo of the olivine are scaled by intensity of red and green respectively to show the diverse compositions in OK08, and where they are likely placed in the entire plumbing architecture. Stars represent seismicity during the period.